# CS205b/CME306 

Lecture 18

The full Navier-Stokes equations are

$$
\begin{aligned}
\rho_{t}+\nabla \cdot(\rho \mathbf{u}) & =0 \\
(\rho \mathbf{u})_{t}+\nabla \cdot\left(\rho \mathbf{u u ^ { T }}+p \mathbf{I}\right) & =\nabla \cdot \boldsymbol{\tau}+\rho \mathbf{g} \\
E_{t}+\nabla \cdot((E+p) \mathbf{u}) & =\nabla \cdot(\boldsymbol{\tau} \mathbf{u})+\nabla \cdot(k \nabla T)
\end{aligned}
$$

where $T$ is the temperature, $k$ is the thermal conductivity, and

$$
\begin{aligned}
\boldsymbol{\tau} & =\left(\begin{array}{cc}
2 \mu u_{x}+\lambda\left(u_{x}+v_{y}\right) & \mu\left(u_{y}+v_{x}\right) \\
\mu\left(u_{y}+v_{x}\right) & 2 \mu v_{y}+\lambda\left(u_{x}+v_{y}\right)
\end{array}\right) \\
& =\mu\binom{\nabla u}{\nabla v}+\mu\binom{\nabla u}{\nabla v}^{T}+\lambda\left(u_{x}+v_{y}\right) \mathbf{I} \\
& =\mu\left(\nabla \mathbf{u}+\nabla \mathbf{u}^{T}\right)+\lambda(\nabla \cdot \mathbf{u}) \mathbf{I} .
\end{aligned}
$$

The parameter $\lambda$ is often chosen to make the $\nabla \cdot \boldsymbol{\tau}=0$. The latter criterion is called Stokes Hypothesis and results in $\lambda=-\frac{2}{3} \mu$ in 3D and $\lambda=-\mu$ in 2D. The Navier-Stokes equations simplify under the incompressibility assumption to

$$
\begin{align*}
\nabla \cdot \mathbf{u} & =0 \\
\rho_{t}+\mathbf{u} \cdot \nabla \rho & =0 \\
\mathbf{u}_{t}+\mathbf{u} \cdot \nabla \mathbf{u}+\frac{\nabla p}{\rho} & =\frac{\nabla \cdot \boldsymbol{\tau}}{\rho}+\mathbf{g} \\
e_{t}+\mathbf{u} \cdot \nabla e & =\frac{\operatorname{tr}(\boldsymbol{\tau} \nabla \mathbf{u})}{\rho}+\frac{\nabla \cdot(k \nabla T)}{\rho}-\mathbf{u} \cdot \mathbf{g} \tag{1}
\end{align*}
$$

where $\boldsymbol{\tau}$ simplifies to

$$
\boldsymbol{\tau}=\mu\left(\begin{array}{cc}
2 u_{x} & u_{y}+v_{x} \\
u_{y}+v_{x} & 2 v_{y}
\end{array}\right)=\mu\binom{\nabla u}{\nabla v}+\mu\binom{\nabla u}{\nabla v}^{T}=\mu\left(\nabla \mathbf{u}+\nabla \mathbf{u}^{T}\right) .
$$

## 1 Heat Equation

By removing the viscosity and forcing terms from equation 1 one has

$$
\begin{equation*}
e_{t}+\mathbf{u} \cdot \nabla e=\frac{\nabla \cdot(k \nabla T)}{\rho} \tag{2}
\end{equation*}
$$

The assumptions that $e$ and $T$ satisfy the relationship

$$
d e=c_{v} d T
$$

simplifies equation 2 to

$$
c_{v} T_{t}+c_{v} \mathbf{u} \cdot \nabla T=\frac{\nabla \cdot(k \nabla T)}{\rho}
$$

which can be further simplified to the standard heat equation

$$
\begin{equation*}
T_{t}=\frac{1}{\rho c_{v}} \nabla \cdot(k \nabla T) \tag{3}
\end{equation*}
$$

by ignoring the effects of convection, i.e. setting $\mathbf{u}=0$. (Note that the assumption $\mathbf{u}=0$ will also eliminate the viscosity and forcing terms from the energy equation.) If $k$ is constant, this can be written as

$$
T_{t}=\frac{k}{\rho c_{v}} \Delta T .
$$

Applying explicit Euler time discretization to equation 3 results in

$$
\frac{T^{n+1}-T^{n}}{\Delta t}=\frac{1}{\rho c_{v}} \nabla \cdot\left(k \nabla T^{n}\right)
$$

where either Dirichlet or Neumann boundary conditions can be applied on the boundaries of the computational domain. Assuming that $\rho$ and $c_{v}$ are constants allows us to rewrite this equation as

$$
\frac{T^{n+1}-T^{n}}{\Delta t}=\nabla \cdot\left(\hat{k} \nabla T^{n}\right)
$$

with $\hat{k}=\frac{k}{\rho c_{v}}$. Standard central differencing (second order accurate) can be used for the spatial derivatives as in

$$
\frac{\hat{k}_{i+\frac{1}{2}, j}\left(\frac{T_{i+1, j}-T_{i, j}}{\Delta x}\right)-\hat{k}_{i-\frac{1}{2}, j}\left(\frac{T_{i, j}-T_{i-1, j}}{\Delta x}\right)}{\Delta x}
$$

A time step restriction of

$$
\Delta t \hat{k}\left(\frac{2}{\Delta x^{2}}+\frac{2}{\Delta y^{2}}+\frac{2}{\Delta z^{2}}\right) \leq 1
$$

is needed for stability. If we $\Delta x=\Delta y$, then this is

$$
2 n \frac{\Delta t}{\Delta x^{2}} \hat{k} \leq 1
$$

where $n$ is the dimension ( $n=2$ in 2 D and $n=3$ in 3 D ).
Implicit Euler time discretization

$$
\begin{equation*}
\frac{T^{n+1}-T^{n}}{\Delta t}=\nabla \cdot\left(\hat{k} \nabla T^{n+1}\right) \tag{4}
\end{equation*}
$$

avoids this time step stability restriction. This equation can be rewritten as

$$
\begin{equation*}
T^{n+1}-\Delta t \nabla \cdot\left(\hat{k} \nabla T^{n+1}\right)=T^{n} \tag{5}
\end{equation*}
$$

discretizing the $\nabla \cdot\left(\hat{k} \nabla T^{n+1}\right)$ term using central differencing. For each unknown, $T_{i}^{n+1}$, equation 5 is used to fill in one row of a matrix creating a linear system of equations. Since the resulting matrix is symmetric, a number of fast linear solvers can be used (e.g. a PCG method with an incomplete Choleski preconditioner, see Golub and Van Loan [1]). Equation 4 is first order accurate in time and second order accurate in space, and $\Delta t$ needs to be chosen proportional to $\Delta x^{2}$ in order to obtain an overall asymptotic accuracy of $O\left(\Delta x^{2}\right)$. However, the stability of the implicit Euler method allows one to chose $\Delta t$ proportional to $\Delta x$ saving dramatically on CPU time. The Crank-Nicolson scheme

$$
\frac{T^{n+1}-T^{n}}{\Delta t}=\frac{1}{2} \nabla \cdot\left(\hat{k} \nabla T^{n+1}\right)+\frac{1}{2} \nabla \cdot\left(\hat{k} \nabla T^{n}\right)
$$

can be used to achieve second order accuracy in both space and time with $\Delta t$ proportional to $\Delta x$. For the Crank-Nicolson scheme,

$$
T^{n+1}-\frac{\Delta t}{2} \nabla \cdot\left(\hat{k} \nabla T^{n+1}\right)=T^{n}+\frac{\Delta t}{2} \nabla \cdot\left(\hat{k} \nabla T^{n}\right)
$$

is used to create a symmetric linear system of equations for the unknowns $T_{i}^{n+1}$. Again, all spatial derivatives are computed using standard central differencing.

Why not always use Crank-Nicholson, as it gives second order accuracy and no time step restriction? Let us look at the solution as $\Delta t \rightarrow \infty$. Backward Euler gives

$$
\Delta T^{n}=0
$$

which is the correct steady state solution. Crank-Nicholson gives

$$
\Delta T^{n+1}=-\Delta T^{n}
$$

In $1 D$ this is

$$
T_{x x}^{n+1}=-T_{x x}^{n}
$$

This shows that the curvature is changing sign at each time step. So the problem with CrankNicholson is that as $\Delta t$ gets very large, you get oscillations, whereas with backward Euler, you get the steady-state solution.

## References

[1] Golub, G. and Van Loan, C., Matrix Computations, The Johns Hopkins University Press, Baltimore, 1989.

